

# History of the Universe (a current view)



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# The Big Bang theory

- Discoveries in astronomy and physics suggest that the Universe had a definite beginning.
- The Big Bang theory is the prevailing theory that describes the origin and evolution of the Universe.
- It is considered to be our best theory of cosmology because it explains most experimental observations.



# What was the Big Bang?

- About 13.7 billion years ago, the entire Universe was compressed into a **singularity** – a place with zero volume and infinite density.
- Matter, energy, space and time all began inside the singularity.
- At the moment of the Big Bang, the Universe started to expand, cool and become less dense.
- The Universe is still expanding today.

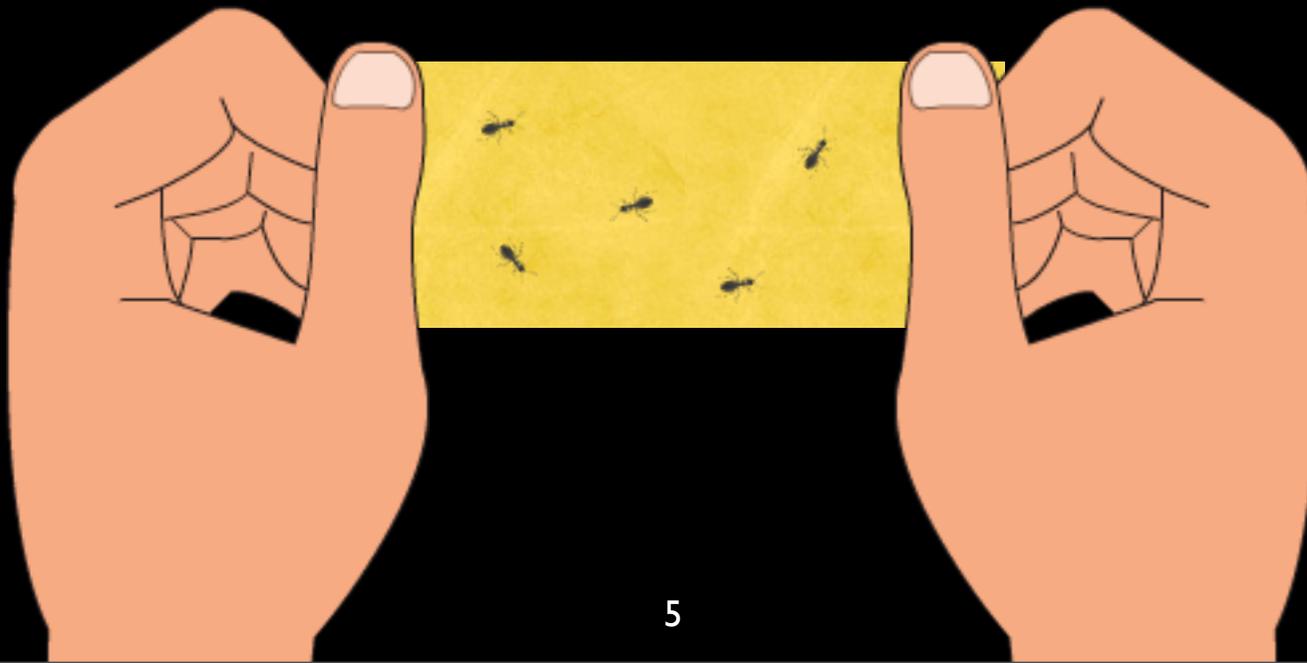
# What the Big Bang wasn't

- The Big Bang wasn't an explosion that happened somewhere in space.
- The Universe didn't appear somewhere in space.
- The Big Bang created space and time.



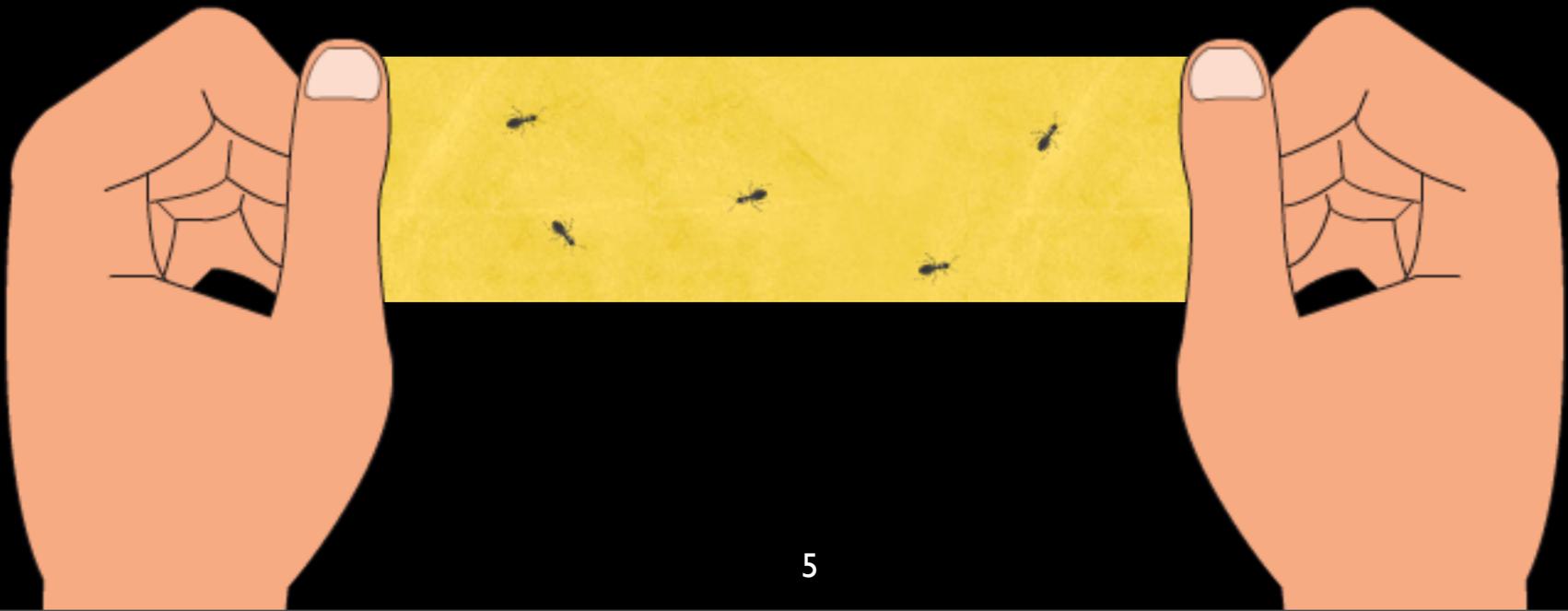
# What is an expanding Universe?

- Galaxies move apart because space itself is expanding.
- They are not moving apart because of some past explosion.

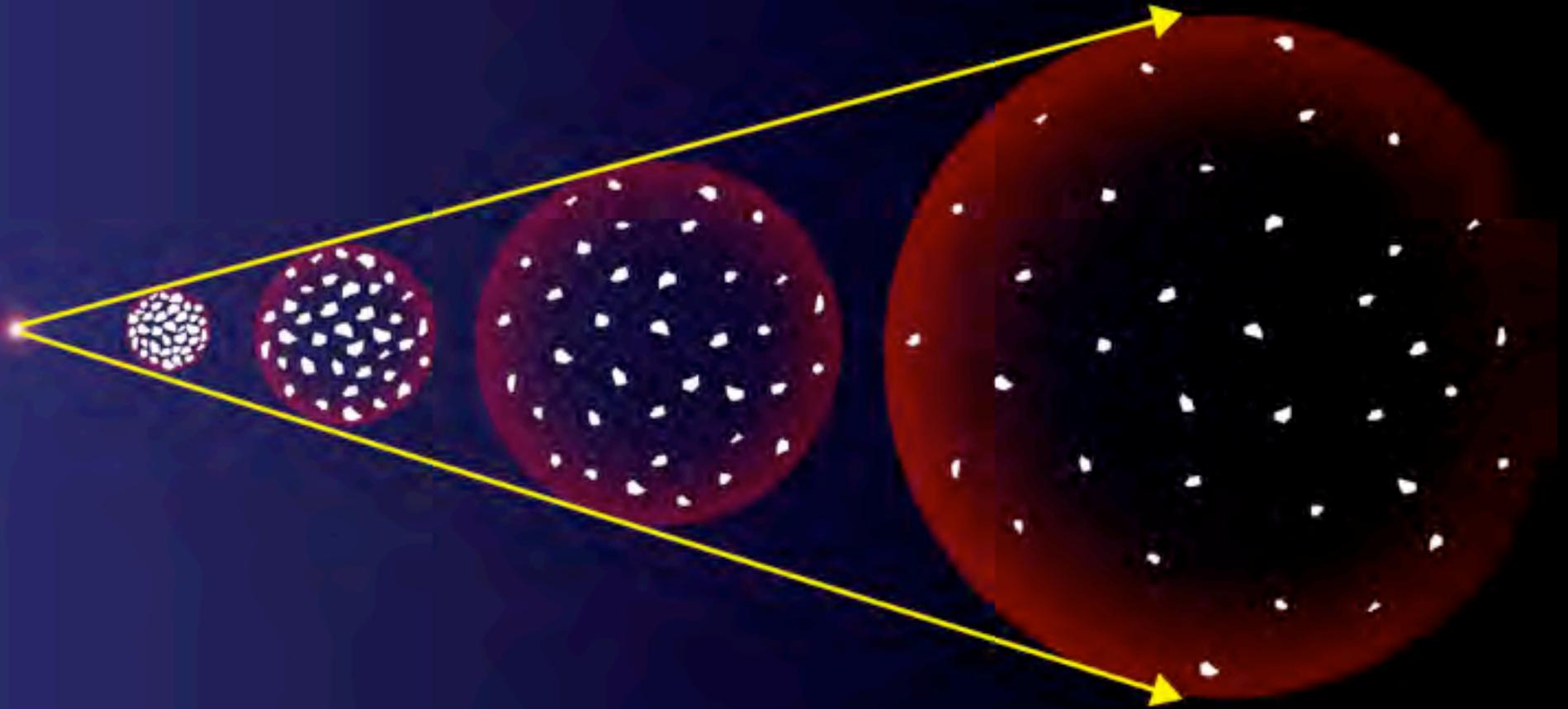


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# Expansion of the Universe



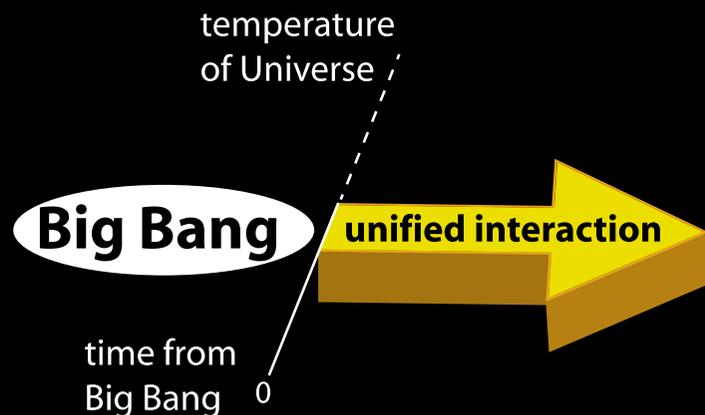
Galaxies move apart as the Universe expands.

# The beginning of the Universe

- Big Bang theory doesn't attempt to explain why the Universe was created, or what (if anything) might have existed before it.
- Our ideas about the very early Universe are, at best, speculative.
- The first  $10^{-43}$  seconds of the Universe is called the Planck era. Conditions were so extreme that we suspect quantum behaviour was dominant, including quantum gravity.
- We don't yet have a theory of quantum gravity, but we do have theories that explain what happened from this time on.

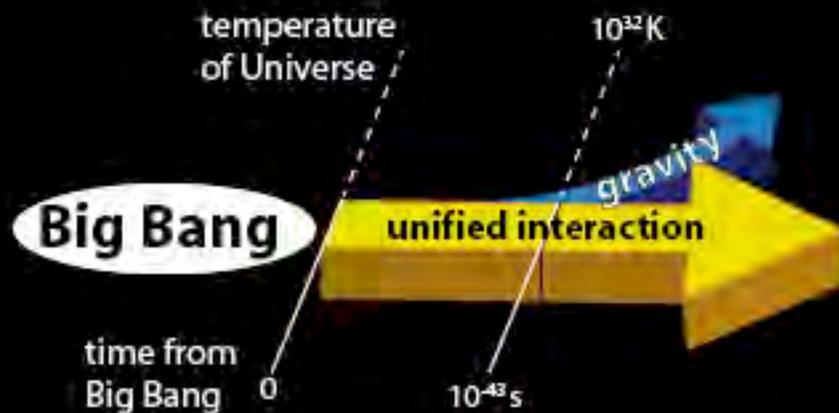
# Fundamental interactions

- All forces in the Universe can be attributed to four fundamental interactions between particles.
- Physicists believe that all known interactions were unified at the moment of the Big Bang.
- Shortly after the Big Bang, this unified interaction began to separate into gravitational, strong, weak and electromagnetic interactions.



# Fundamental interactions

- $10^{-43}$  seconds after the Big Bang, gravity separated from the unified interaction.



$10^{32}$ K

$10^{40}$

$10^{30}$

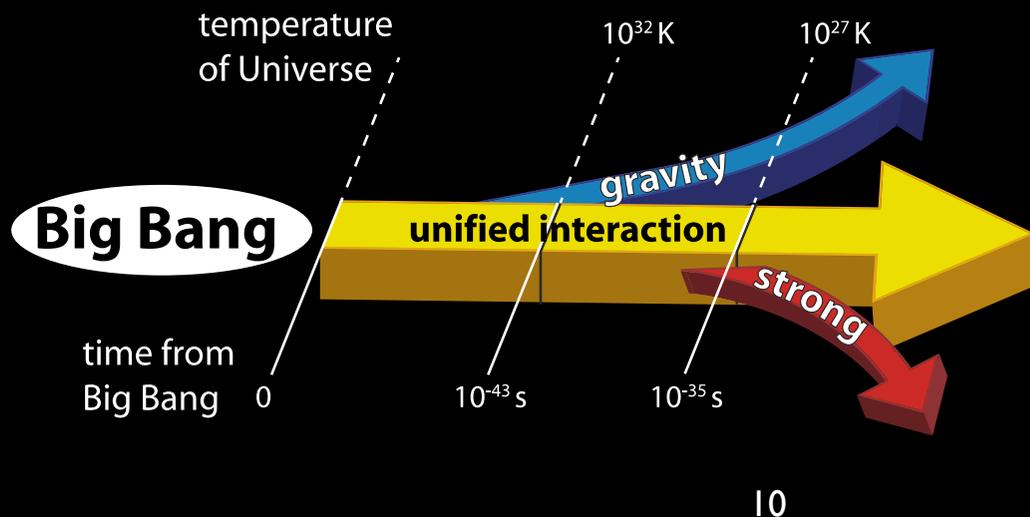
$10^{20}$

$10^{10}$

1

# Fundamental interactions

- $10^{-35}$  seconds after the Big Bang, the strong interaction separated.
- This released a vast amount of energy, making the Universe expand at an extraordinary rate.



$10^{27}$ K

$10^{40}$

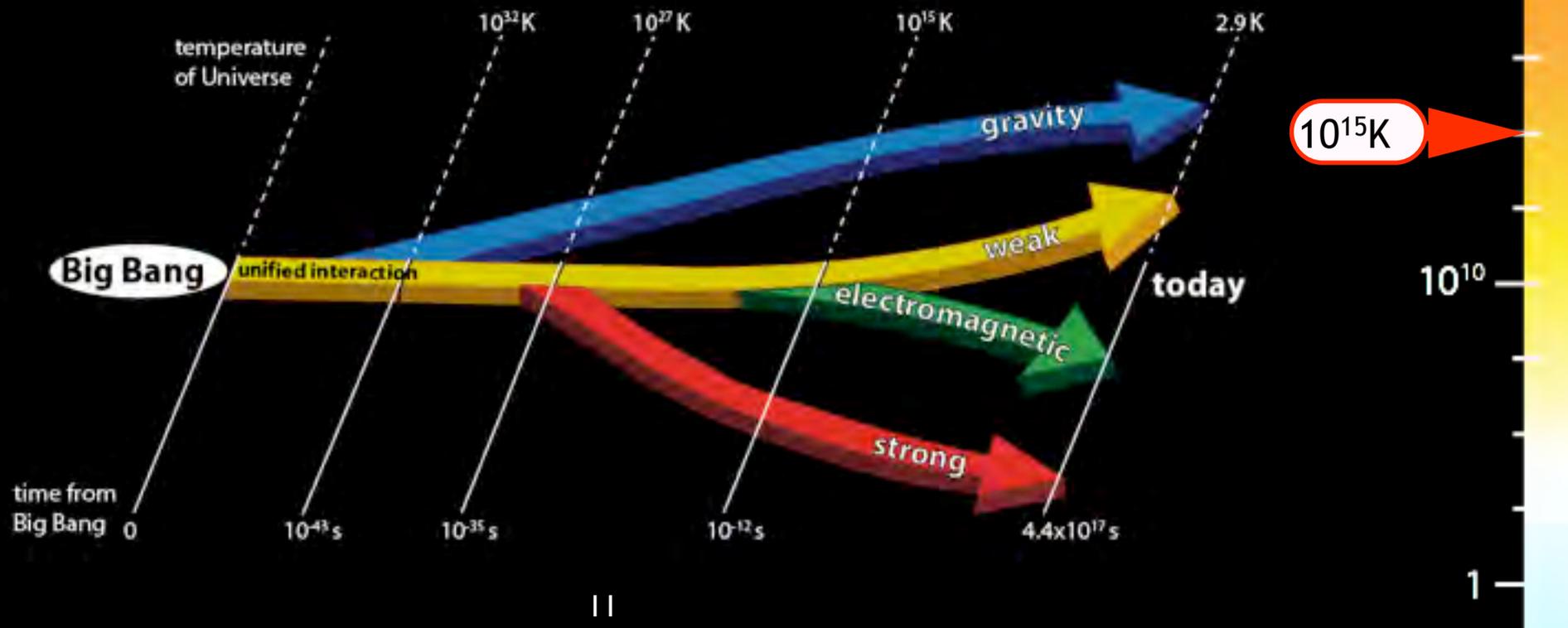
$10^{30}$

$10^{20}$

$10^{10}$

# Fundamental interactions

- $10^{-12}$  seconds after the Big Bang, the electromagnetic and the weak interactions separated. The four fundamental interactions were now distinct, as they remain to this day.

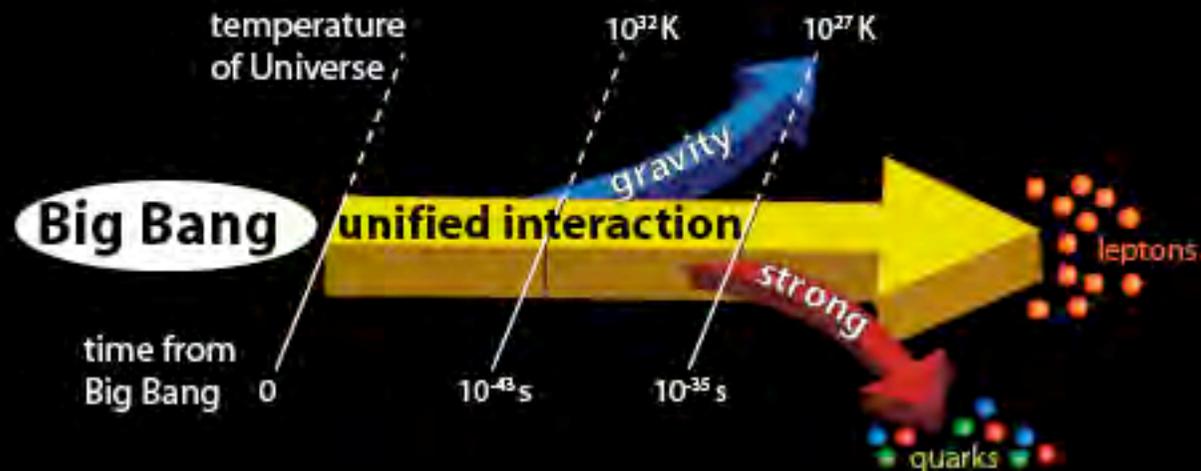


# The evolution of matter

quark era	$10^{-12}$ – $10^{-6}$ seconds after the Big Bang
hadron era	$10^{-6}$ – 1 second after the Big Bang
lepton era	1 sec – 10 seconds after the Big Bang
radiation era	10 sec – 300 000 years after the Big Bang

# The **quark** era ( $10^{-12}$ – $10^{-6}$ s)

- When the strong interaction separated from the unified interaction, matter separated into quarks and leptons.
- Quarks 'feel' the strong interaction, but leptons don't.
- The temperature of the Universe was too high for quarks to combine.



# The **hadron** era ( $10^{-6} - 1$ s)

- As the temperature of the Universe fell, quarks combined to form hadrons: protons, neutrons and their antiparticles.
- Almost equal quantities of particles and antiparticles were created.
- As the temperature fell further, most particles annihilated their antiparticles.
- A small excess of particles was left, which accounts for all hadrons in the Universe today.

proton



neutron

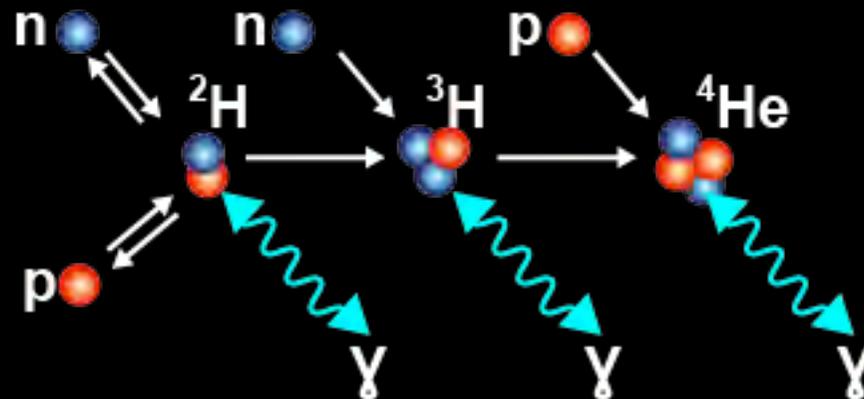
# The **lepton** era (1 s – 10 s)

- Once most hadrons had been annihilated, leptons became the dominant form of matter: electrons, neutrinos and their antiparticles.
- As the Universe continued to cool, most leptons and antileptons annihilated each other.
- A small excess of leptons was left, which accounts for all the leptons in the Universe today.



# The radiation era (from 10 s)

- Particle-antiparticle interactions left the Universe full of radiation (photons).
- Protons and neutrons combined to form hydrogen and helium nuclei in a process called nucleosynthesis.
- Photons were continually absorbed and emitted by particles, so the Universe was opaque to radiation.



# The radiation era

- Towards the end of the radiation era, 300 000 years after the Big Bang, hydrogen and helium nuclei started to capture electrons to form stable atoms.
- The density of the Universe continued to decrease as it expanded.
- The Universe became transparent to light about 370 000 years after the Big Bang.

# The formation of stars

- Over the next 100 million years or so, the Universe continued to expand and cool.
- Gravitational variations caused matter to clump together and become hotter and denser.
- 100–200 million years after the Big Bang, matter began to coalesce and form stars.
- Fusion of light elements in stars released energy and formed elements from carbon to iron.

‘Eagle nebula pillars’ from NASA/ESA  
[http://commons.wikimedia.org/wiki/File:Eagle\\_nebula\\_pillars\\_complete.jpg](http://commons.wikimedia.org/wiki/File:Eagle_nebula_pillars_complete.jpg)

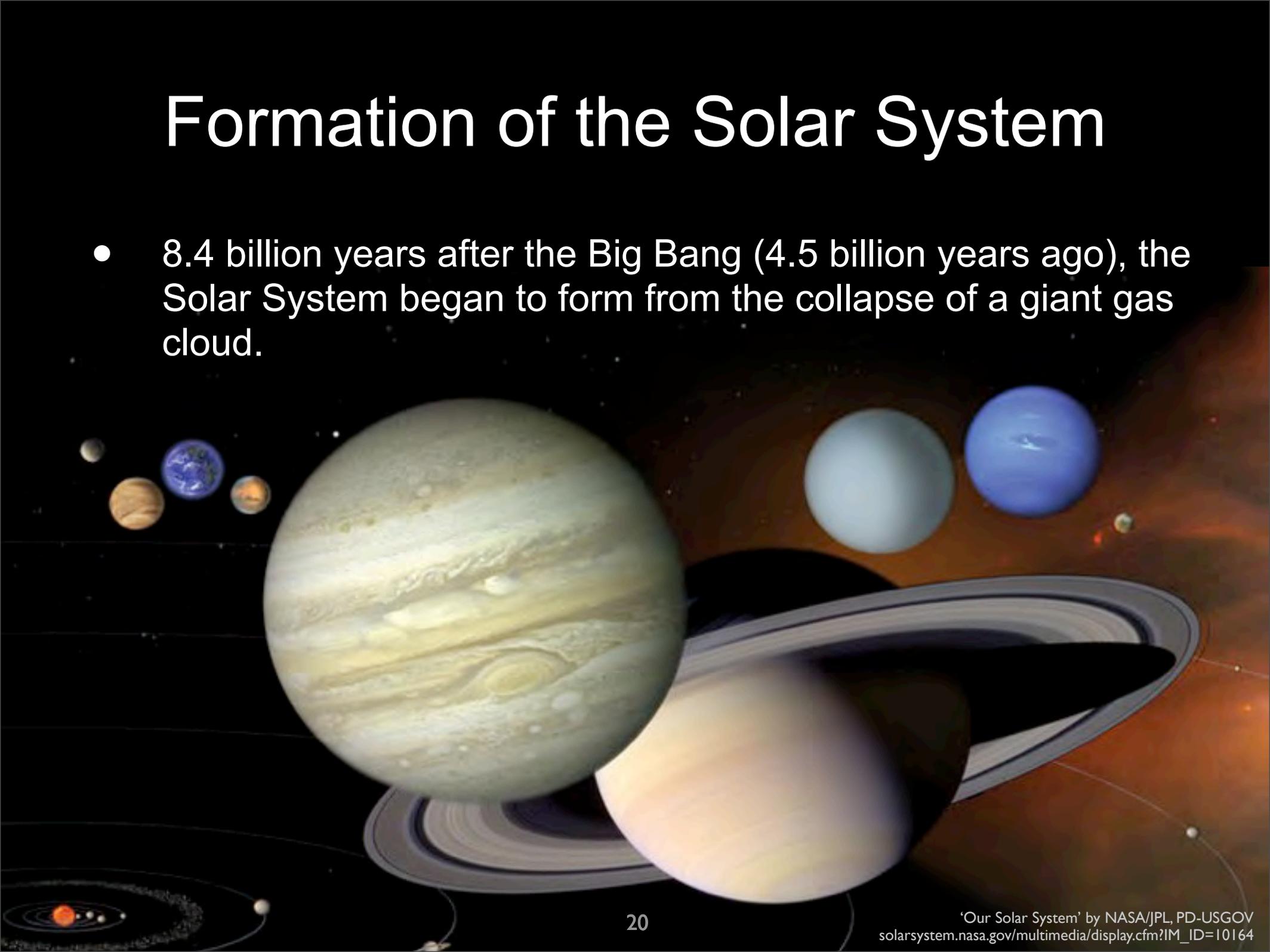
# The formation of galaxies

- A billion years after the Big Bang, stars grouped to form galaxies.
- Galaxies formed clusters.



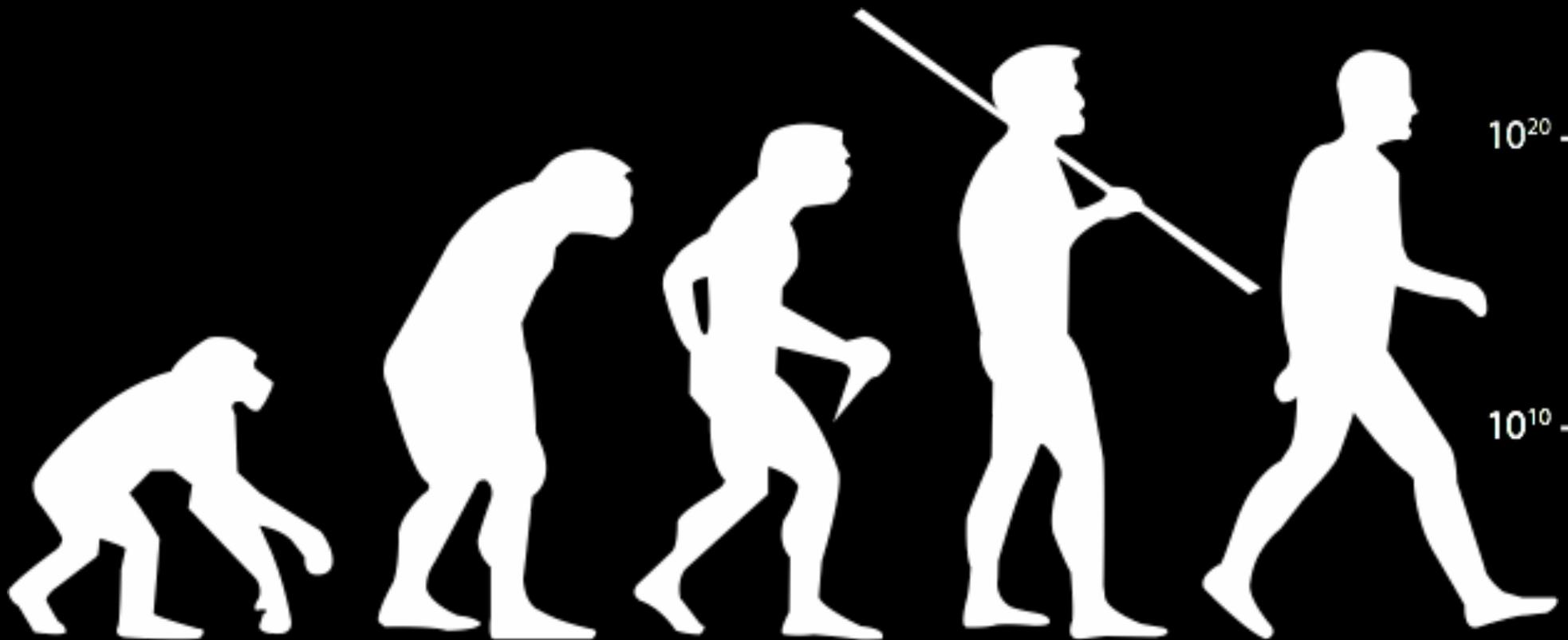
# Formation of the Solar System

- 8.4 billion years after the Big Bang (4.5 billion years ago), the Solar System began to form from the collapse of a giant gas cloud.



# Humans appear on Earth

- 13.7 billion years after the Big Bang, human life began on Earth.

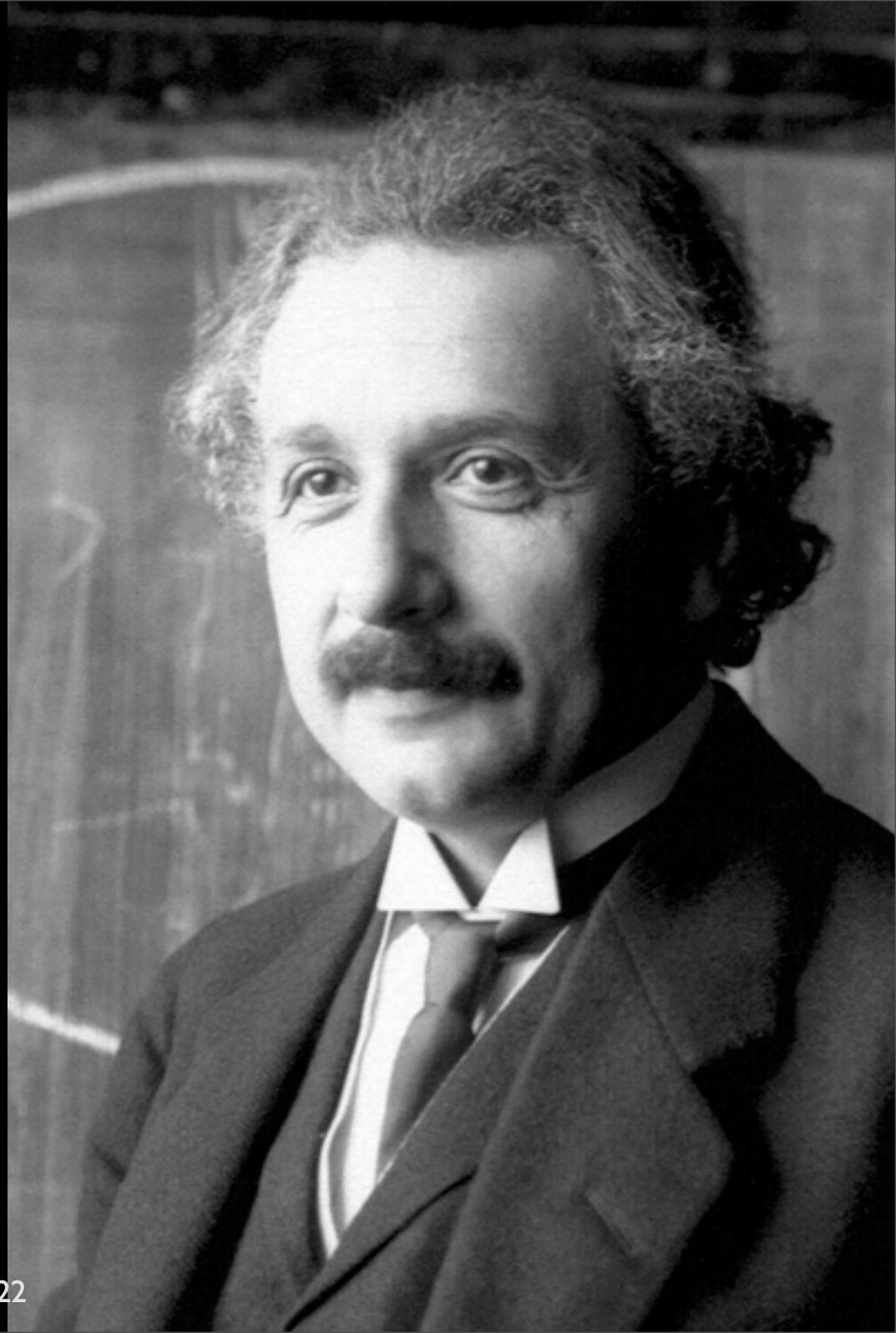


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"The most incomprehensible thing about the Universe is that it is comprehensible."

Albert Einstein, 1935

photo by Ferdinand Schmutzer,  
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